Progress on Light-Ion Fusion Reactions with Three-Nucleon Forces

22nd European Conference on Few Body Problems in Physics

Cracow, September 12th 2013.

Lawrence Livermore National Laboratory

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LLNL-PRES-643689

This work was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under contract DE-AC52-07NA27344. Lawrence Livermore National Security, LLC



Guillaume Hupin

Some words about the ingredients of an ab initio calculation

A high precision nuclear Hamiltonian



- The nuclear interaction has a strong repulsive core.
- This makes nuclear structure calculations to converge slowly.

...and also NNN interaction



- NNN interaction effects are important.
- This is ~100 times numerically costlier.



Ab initio NCSM/RGM: formalism for binary clusters

S. Quaglioni and P. Navrátil, Phys. Rev. Lett. 101, 092501 (2008); Phys. Rev. C 79, 044606 (2009)



Schrödinger equation on channel basis:

RGM accounts for: 1) interaction (Hamiltonian kernel), 2) Pauli principle (Norm kernel) between clusters and NCSM accounts for: internal structure of clusters



Going around the hard core problem

E. Jurgenson, Navrátil, R. J. Furnstahl Phys. Rev. Lett. 103 (2009)

In configuration interaction methods we need to soften interaction to address the hard core We use the Similarity-Renormalization-Group (SRG) method





Flow parameter







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Including the NNN force into the NCSM/RGM approach nucleon-nucleus formalism

$$\left\langle \Phi_{\nu'r'}^{J^{\pi}T} \left| \hat{A}_{\nu'} V^{NNN} \hat{A}_{\nu} \right| \Phi_{\nu r}^{J^{\pi}T} \right\rangle = \left\langle \begin{array}{c} \begin{pmatrix} (A-1) \\ r' \end{pmatrix} \\ r' \end{pmatrix} \left| \begin{array}{c} (A-1) \\ (a'=1) \end{pmatrix} \right| \begin{pmatrix} (A-1) \\ (a'=1) \end{pmatrix} \\ (a''=1) \end{pmatrix} \left| \begin{array}{c} (A-1) \\ (a''=1) \end{pmatrix} \right| \begin{pmatrix} (A-1) \\ (a''=1) \end{pmatrix} \\ (a''') \end{pmatrix} \right\rangle$$

$$\mathcal{V}_{\nu'\nu}^{NNN}(r,r') = \sum R_{n'l'}(r')R_{nl}(r) \left[\frac{(A-1)(A-2)}{2} \left\langle \Phi_{\nu'n'}^{J^{\pi}T} | V_{A-2A-1A}(1-2P_{A-1A}) | \Phi_{\nu n}^{J^{\pi}T} \right\rangle \right]$$

$$\overset{(a)}{=} \frac{(A-1)(A-2)(A-3)}{2} \left\langle \Phi_{\alpha_{1}'}^{J^{\pi}T} | P_{A-1A}V_{A-3A-2A-1} | \Phi_{\nu n}^{J^{\pi}T} \right\rangle \left[\cdot \frac{(A-1)(A-2)(A-3)}{2} \left\langle \Phi_{\nu'n'}^{J^{\pi}T} | P_{A-1A}V_{A-3A-2A-1} | \Phi_{\nu n}^{J^{\pi}T} \right\rangle \right] \cdot \frac{(A-1)(A-2)(A-3)}{2} \left\langle \Psi_{\alpha_{1}'}^{(A-1)} | a_{h}^{+}a_{i}^{+}a_{j}^{-}a_{m}a_{l}a_{k} | \Psi_{\alpha_{1}}^{(A-1)} \right\rangle_{SD}$$



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n-⁴He scattering: study of the model space convergence





- Calculations are done using the chiral EFT interaction at N³LO for the NN and N²LO for the NNN.
- SRG flow parameter is λ =2.0 fm⁻¹
- Phase-shifts are to a good extent independent from ħΩ.



n-⁴He scattering: study of the HO model-space convergence





- Convergence pattern is good.
- In the following, we adopt ħΩ=20 MeV.



n-⁴He scattering: study of the RGM convergence in the NNN case

G. Hupin, J. Langhammer et al. ArXiv:1308.2700





• The convergence of the phase-shifts is reached when the first six excited states of ⁴He are included.



n-⁴He scattering: study of the RGM convergence in the NN case

P. Navrátil and S. Quaglioni, PRC83 044609 (2011)





Study of the SRG flow parameter dependence

G. Hupin, J. Langhammer et al. ArXiv:1308.2700



Unitarity behavior of the calculated phase-shifts 120 ${}^{2}P_{3/2}$ ${}^{2}P_{3/2}$ ${}^{2}P_{3/2}$ 90 ${}^{2}P_{1/2}$ ${}^{2}P_{1/2}$ ${}^{2}P_{1/2}$ 60 30[deg] ${}^{2}D_{3/2}$ ${}^{2}D_{3}{}_{/2}$ $^{2}D_{3/2}$ (b) NN+3N-induced (a) NN-only (c) NN+3N-full 0 \sim -30 $^{2}S_{1/2}$ ${}^{2}S_{1/2}$ ${}^{2}S_{1/2}$ $\lambda = 1.88 \text{ fm}^{-1}$ $\lambda = 2.0 \text{ fm}^{-1}$ $\lambda = 1.88 \text{ fm}^{-1}$ - $\lambda = 2.0 \text{ fm}^{-1}$ $\lambda = 1.88 \text{ fm}^{-1}$ -60 $-\lambda = 2.0 \text{ fm}$ $\lambda = 2.24 \text{ fm}^{-1}$ -90 n^{-4} He (g.s., 0⁺, 0⁻, 2⁻, 2⁻, 1⁻, 1⁻) n^{-4} He (g.s., 0⁺, 0⁻, 2⁻, 2⁻, 1⁻, 1⁻) n^{-4} He (g.s., 0⁺, 0⁻, 2⁻, 2⁻) 1212128 8 4 8 0 4 0 4 160 E_{kin} [MeV] E_{kin} [MeV] E_{kin} [MeV] Comparison at Nmax=13 between NN-only, NN+NNN-ind and NN+NNN between λ =2.0 and 1.88 fm⁻¹

• The SRG evolution is to a good extend unitary.



n-⁴He scattering: NN versus NNN interactions

G. Hupin, J. Langhammer et al. ArXiv:1308.2700





- The NNN interactions affect mostly the P waves.
- The largest spin-orbit splitting between *P* waves is obtained with NN+NNN.
- The agreement of the P_{3/2} phase-shifts between NN-only and NN+NNN forces is accidental.

Comparison between NN+NNN -ind and NN+NNN at Nmax=13 with six ⁴He states.



n-⁴He scattering: NCSM/RGM versus experiment

G. Hupin, J. Langhammer et al. ArXiv:1308.2700







only

p-⁴He scattering: NCSM/RGM versus experiment







Analyzing power and differential cross section

G. Hupin, J. Langhammer et al. ArXiv:1308.2700





- Good overall reproduction of the experimental data yet it cannot fully explain the A_y puzzle.
- So far we have to restrict the number of partial waves in the calculation due to the model-space truncation related to NNN interactions.

Comparison between NN-only, NN+NNN –ind, NN+NNN and experiment of the differential cross section at $E_n=17.6$ MeV and analyzing power at $E_n=15$ MeV.



Analyzing power and differential cross section





Including the NNN force into the NCSM/RGM approach deuteron-nucleus formalism

$$\left\langle \Phi_{\nu\nu'}^{r\pi T} \left| \hat{A}_{\nu} V^{NNN} \hat{A}_{\nu} \right| \Phi_{\nu T}^{r\pi T} \right\rangle = \left\langle \underbrace{(A-2)}_{r(a=2)} \right\rangle V^{NNN} \left(1 - \sum_{i=1}^{A-2} \sum_{k=A-1}^{A} \hat{P}_{i,k} + \sum_{i< j=1}^{A-2} \hat{P}_{i,A-1} \hat{P}_{j,A} \right) \left| \underbrace{(A-2)}_{(a=2)} \right\rangle$$
Direct
$$\left\{ \underbrace{(A-2)}_{\nu\nu'} \left| \hat{A}_{\nu} V^{NNN} \hat{A}_{\nu} \right| \Phi_{\nu T}^{r} \right\} = \left\langle \underbrace{(A-2)}_{r(a=2)} \right\rangle V^{NNN} \left(1 - \sum_{i=1}^{A-2} \sum_{k=A-1}^{A} \hat{P}_{i,k} + \sum_{i< j=1}^{A-2} \hat{P}_{i,A-1} \hat{P}_{j,A} \right) \left| \underbrace{(A-2)}_{(a=2)} \right\rangle$$
Exchange
$$\left\{ \underbrace{(A-2)}_{(a=2)} \right\} \left(\underbrace{(A-2)}_{(a=2)} \right) \left(\underbrace{($$

⁴He(*d*,*d*)⁴He: seven pseudo-states





Comparison Expt. R-matrix





Recent progress: coupling of NCSM and NCSM/RGM "NCSMC"

S. Baroni, P. Navrátil and S. Quaglioni PRL110 (2013)

• Methods develop in this presentation to solve the many body problem



• The many body <u>quantum</u> problem best describe by superposition of both

$$\Psi_{NCSMC}^{(A)} = \sum_{\lambda} c_{\lambda} |A\lambda J^{\pi}T\rangle + \sum_{\nu} \int d\vec{r} g_{\nu}(\vec{r}) \hat{A}_{\nu} |\Phi_{\nu\vec{r}}^{(A-a,a)}\rangle$$

NCSMC



NCSMC speeds up convergence of reaction observables

S. Baroni, P. Navrátil and S. Quaglioni PRC87 (2013)



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Conclusions and Outlook



Evolution of stars, birth, main sequence, death

- We are extending the *ab initio* NCSM/RGM approach to describe low-energy reactions with two- and three-nucleon interactions.
- We are able to describe:
 - Nucleon-nucleus collisions with NN+NNN interaction
 - Deuterium-nucleus collisions with NN+NNN interaction
 - NCSMC for single- and twonucleon projectile
- Work in progress
 - Fusion reactions with our best complete *ab initio* approach
 - The present NNN force is "incomplete", need to go to N³LO
 - Scattering of heavier target

